



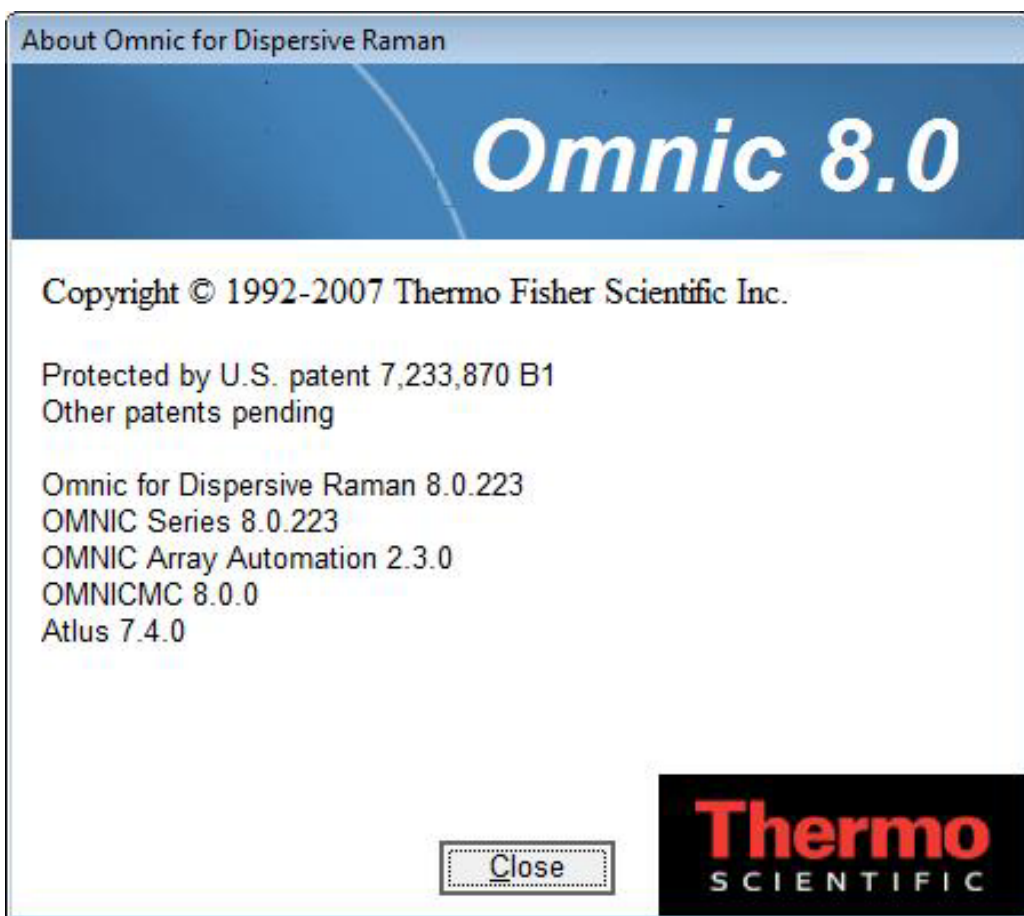
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# OMNIC for DXR Dispersive Raman

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# OMNIC 8.0

- A new version of OMNIC with new features for the DXR and Almega spectrometers is now called OMNIC for Dispersive Raman



# New Features in OMNIC makes Raman Easier



You no longer need to be a Raman expert to produce high quality spectra!

# Experiment Setup Collect Tab

Experiment Setup - C:\My Documents\OMNIC\WRParam\Default.exp

Collect | Bench | Quality | Advanced | CCD Array | Alignment | Mapping | Series

Estimated time for this collection: 00:00:04

Collect exposure time (sec): 1.0

Preview exposure time (sec): 0.5

Sample exposures: 2

Background exposures: 16

Final format: Shifted spectrum (cm-1)

Correction: Fluorescence

Cosmic ray threshold: Medium

☐ Photobleach time (min): 0.0

☒ Preview data collection

☒ Auto exposure: Desired S/N: 100

Maximum collect time (min): 5

File Handling

☐ Save automatically Base name: C:\My Documents\Omic\autosave\0001.spa

Background Handling

☐ Collect background before each sample

☐ Maximum age for background: 1000 minutes

☒ Use smart background

Experiment title: Almaga

Experiment description: Almaga default experiment file

Help Open Save Save As OK Cancel

# Background Correction

- Background correction – refers to correcting for differences in the dark state between pixels on the CCD camera due to the detector dark current (current when there is no light on the element).
- These differences result in a spectrum with subtle patterns in the baseline that have nothing to do with the sample.
- These patterns are easily corrected for by simply collecting an exposure with the shutter closed under the same conditions as the sample collection and subtracting it from the sample measurement.
- The characteristics of this detector background can change slowly over time, so it is not as simple as applying a universal correction.
- OMNIC for Dispersive Raman uses a new approach called Smart Backgrounds.

# Smart Backgrounds

- Fundamentals – runs and saves backgrounds automatically when the system is not being used (OMNIC must remain open).
- Typical collection times are collected and background files stored on the computer hard drive.
- When all data for the smart background model is complete, a message to this effect is displayed. The *Use smart background* option is now enabled in the Background Handling section of the Collect tab in Experiment Setup. (A file named bdata006250.vrb is written to the VRMCa1 folder containing the smart background model.)
- Benefits – saves user time when collecting sample data, can cut user time up to half potentially because the background is already collected



# Autoexposure

- Allows the system to determine best collection time for a sample, makes it easier for a user, less need to be an “expert”.
- Related to Smart Background – autoexposure is only enabled when smart backgrounds are enabled.

The screenshot shows the 'Experiment Setup' dialog box for the file 'C:\My Documents\OMNIC\WRParam\Default.exp'. The 'Advanced' tab is selected. A blue circle highlights the 'Collect exposure time (sec): 1.0', 'Preview exposure time (sec): 0.5', 'Sample exposures: 2', and 'Background exposures: 16' fields. Another blue circle highlights the 'Use smart background' radio button in the 'Background Handling' section. A red circle highlights the 'Auto exposure' checkbox, 'Desired S/N: 100', and 'Maximum collect time (min): 5' fields. Other visible settings include 'Final format: Shifted spectrum (cm-1)', 'Correction: Fluorescence', 'Cosmic ray threshold: Medium', 'Photobleach time (min): 0.0', 'Preview data collection' checked, 'Experiment title: Almega', and 'Experiment description: Almega default experiment file'. The bottom of the dialog has buttons for 'Help', 'Open', 'Save', 'Save As', 'OK', and 'Cancel'.

Experiment Setup - C:\My Documents\OMNIC\WRParam\Default.exp

Collect | Bench | Quality | Advanced | CCD Array | Alignment | Mapping | Series

Estimated time for this collection: 00:00:04

Collect exposure time (sec): 1.0

Preview exposure time (sec): 0.5

Sample exposures: 2

Background exposures: 16

Final format: Shifted spectrum (cm-1)

Correction: Fluorescence

Cosmic ray threshold: Medium

☐ Photobleach time (min): 0.0

☒ Preview data collection

☒ Auto exposure    Desired S/N: 100

Maximum collect time (min): 5

File Handling

☐ Save automatically    Base name: [dropdown]

C:\My Documents\Omnic\autosave\0001.spa

Background Handling

☐ Collect background before each sample

☐ Maximum age for background: 1000 minutes

☒ Use smart background

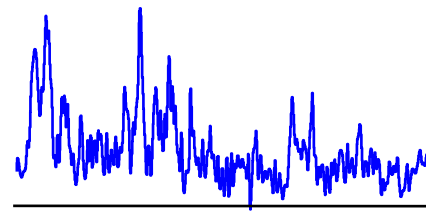
Experiment title: Almega

Experiment description: Almega default experiment file

Help    Open    Save    Save As    OK    Cancel

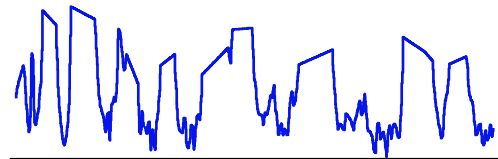
# Setting Exposure Time

## Under Exposed



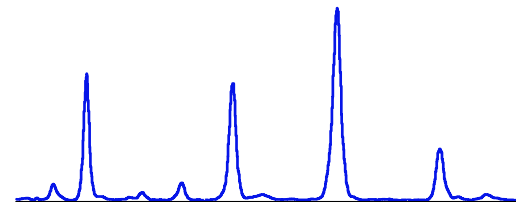
Too short - noisy spectra

## Over Exposed



Too long - detector saturation

## Perfect

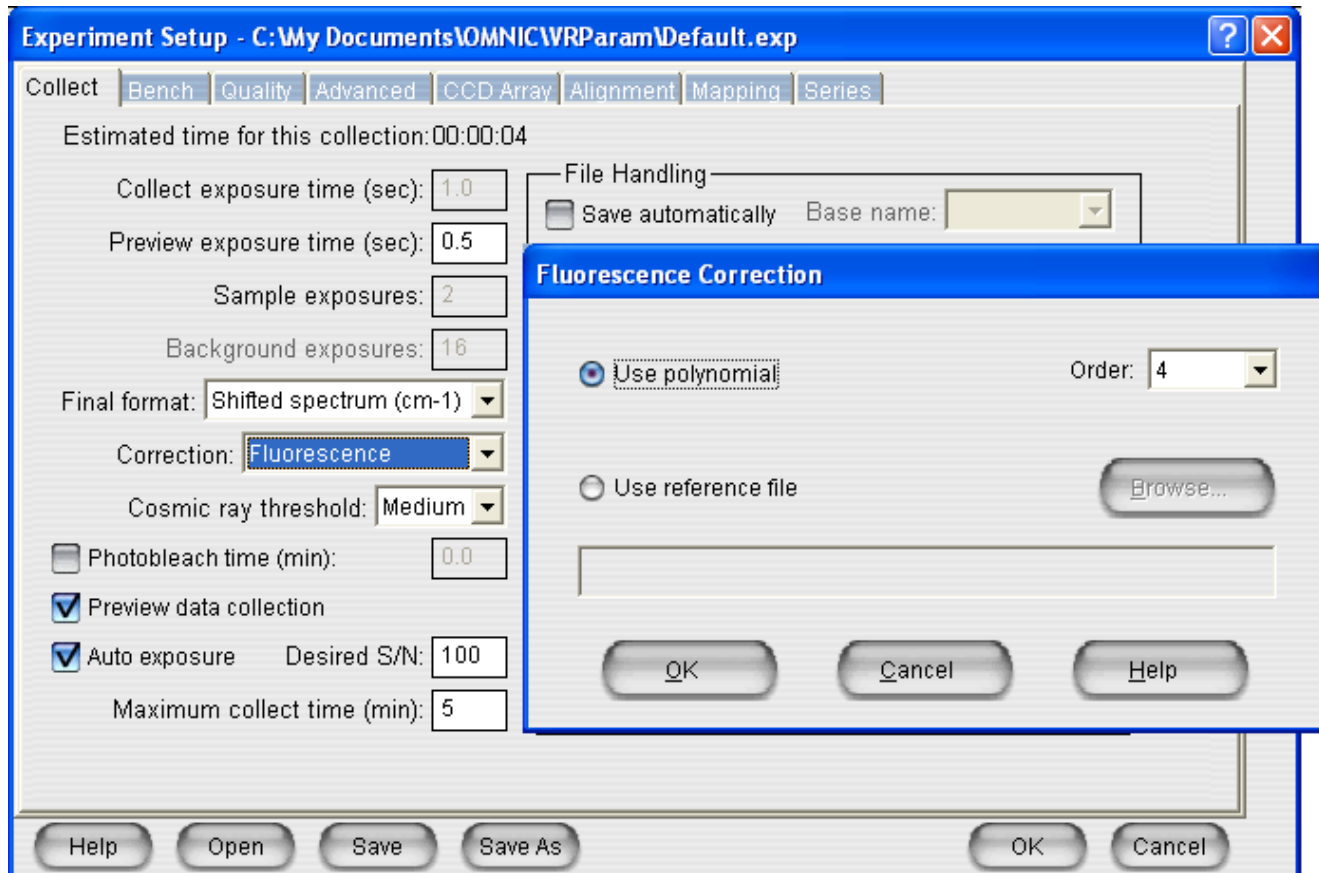


Set correctly – high quality spectra



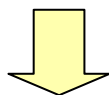
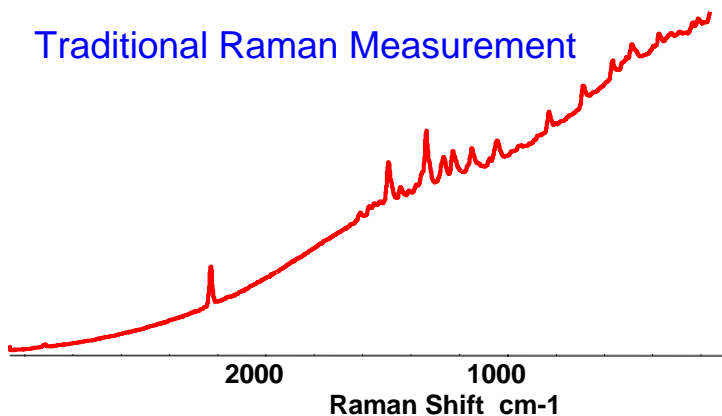
# Autofluorescence Correction

- Some samples may fluoresce, especially with certain lasers. This correction allows the user to remove fluorescence automatically during collection.

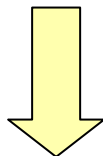


# Automated Fluorescence Rejection

Traditional Raman Measurement



Library Search

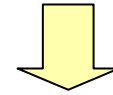
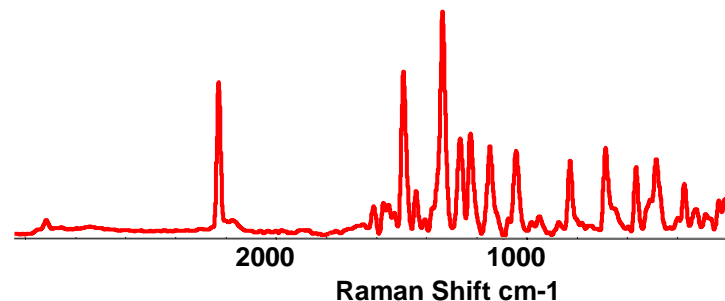


No reasonable matches were found in this search.

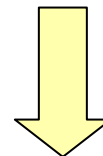
Ragen: 2600 00-450 60

Clipboard View Match List Print Help

Fluorescence Rejection Applied



Library Search



The best match is excellent.  
The match is 5-Methyl-2-  
[(2-nitrophenyl)  
amino]-3-thiophenecarbonitrile

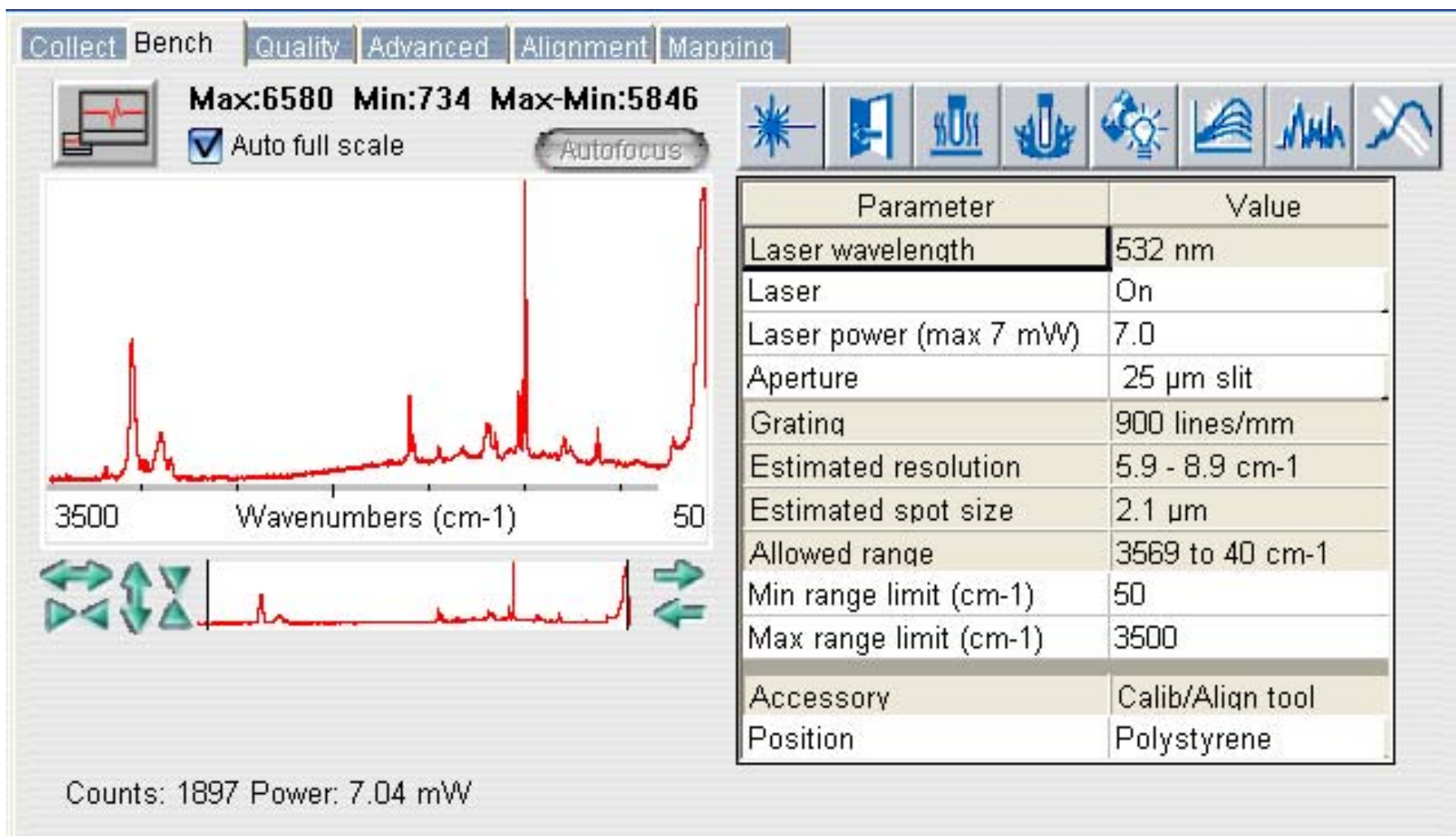
Ragen: 2600 00-450 60

Clipboard View Match List Print Help

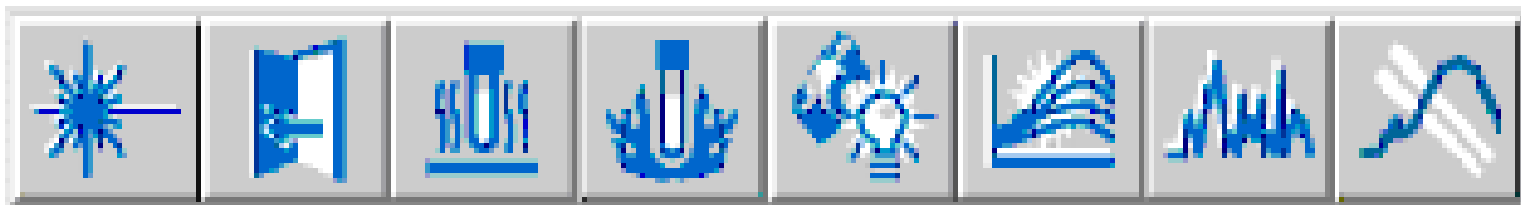
# Automatic Intensity (white light) Correction

- Removes CCD wavelength response, so a spectrum collected at 532 nm can be compared to a spectrum collected at 780 nm. Makes library searching better and more reliable.
- A white light source is used during calibration (a red light is the calibration source when the 780 laser is used). The signal is collected and stored, then collected spectra are corrected after they are collected.
- DXR spectra are always corrected
  - This is done so spectra is reproducible as possible, comparable instrument to instrument and laser to laser, transferable and to remove instrument artifacts.

# Bench Tab



# Status Icons in Bench Tab of Experiment Setup



## Hardware icons

- Laser is off; turn on in OMNIC and turn key clockwise until vertical
- Interlock open (microscope. grating or sample compartment covers are open)

## Icons linked to check marks in Quality tab of Experiment Setup

- Sample may be heating
- Sample is on fire
- CCD overflow (reduce exposure time)
- Photobleaching is occurring (non-zero baseline is decreasing)
- Weak Raman signal (increase aperture and/or exposure time)
- Sample is fluorescing (use photobleaching or purify sample)

# Quality Tab

Collect

Bench

Quality

Advanced

CCD Array

Alignment

Mapping

Series

☒ Use spectral quality checks

Collect Checks

☒ Collect checks

☒ CCD overflow

Spectrum Checks

☒ Spectrum checks

☒ Sample heating☒ Fluorescence☒ Sample burning☒ Photobleaching☒ Weak signal

Minimum:

This corresponds to the Max-Min value above the live display in the Bench tab.

# Advanced Tab

Collect | Bench | Quality | **Advanced** | Alignment | Mapping

Data spacing: 0.984 cm<sup>-1</sup> (2 cm<sup>-1</sup> FT) ☒ Set spacing automatically

Camera temperature: Cooled Laser usage: 790 hours

☒ Laser saver after 300 minutes ☒ Turn laser off when OMNIC closes

Maximum calibration age: 30 days

Maximum alignment age: 30 days ☒ Recalibrate after alignment

Maximum smart background age: 180 days

Macro for Go button:

Autofocus

☐ Before collection ☐ Ignore fluorescence

☐ Autofocus background

☐ Prompt when collecting if laser is off



# Alignment Tab DXR Raman Microscope

- Align tab will only open if align tool is plugged in
- Customer focuses on pinhole, clicks button 2 and Align
- Laser fine adjusts move to maximize photodiode signal

The screenshot shows the 'Alignment' tab of the DXR Raman Microscope software. The interface includes a tab bar at the top with 'Collect', 'Bench', 'Quality', 'Advanced', 'Alignment', and 'Mapping'. The 'Alignment' tab is active, displaying the text 'Auto aligning the 532 nm laser'. On the left, it shows 'Laser signal: 3468' and 'Laser power: 9.8 mW'. On the right, there are three radio buttons for alignment steps: '1. Center pinhole on cross hairs', '2. Align laser to pinhole' (which is selected), and '3. Align spectrograph'. Under '2. Align laser to pinhole', there is a 'Laser Fine Adjust' section with 'Side to side' and 'Up/down' sliders. Under '3. Align spectrograph', there is a 'Spectrograph Fine Adjust' section with 'Side to side' and 'Up/down' sliders. At the bottom, there are two buttons: 'Restore Previous Settings' and 'Cancel'.

Collect Bench Quality Advanced Alignment Mapping

Auto aligning the 532 nm laser

Laser signal: 3468  
Laser power: 9.8 mW

☐ 1. Center pinhole on cross hairs

☒ 2. Align laser to pinhole

Laser Fine Adjust

Side to side: 298

Up/down: 317

☐ 3. Align spectrograph

Spectrograph Fine Adjust

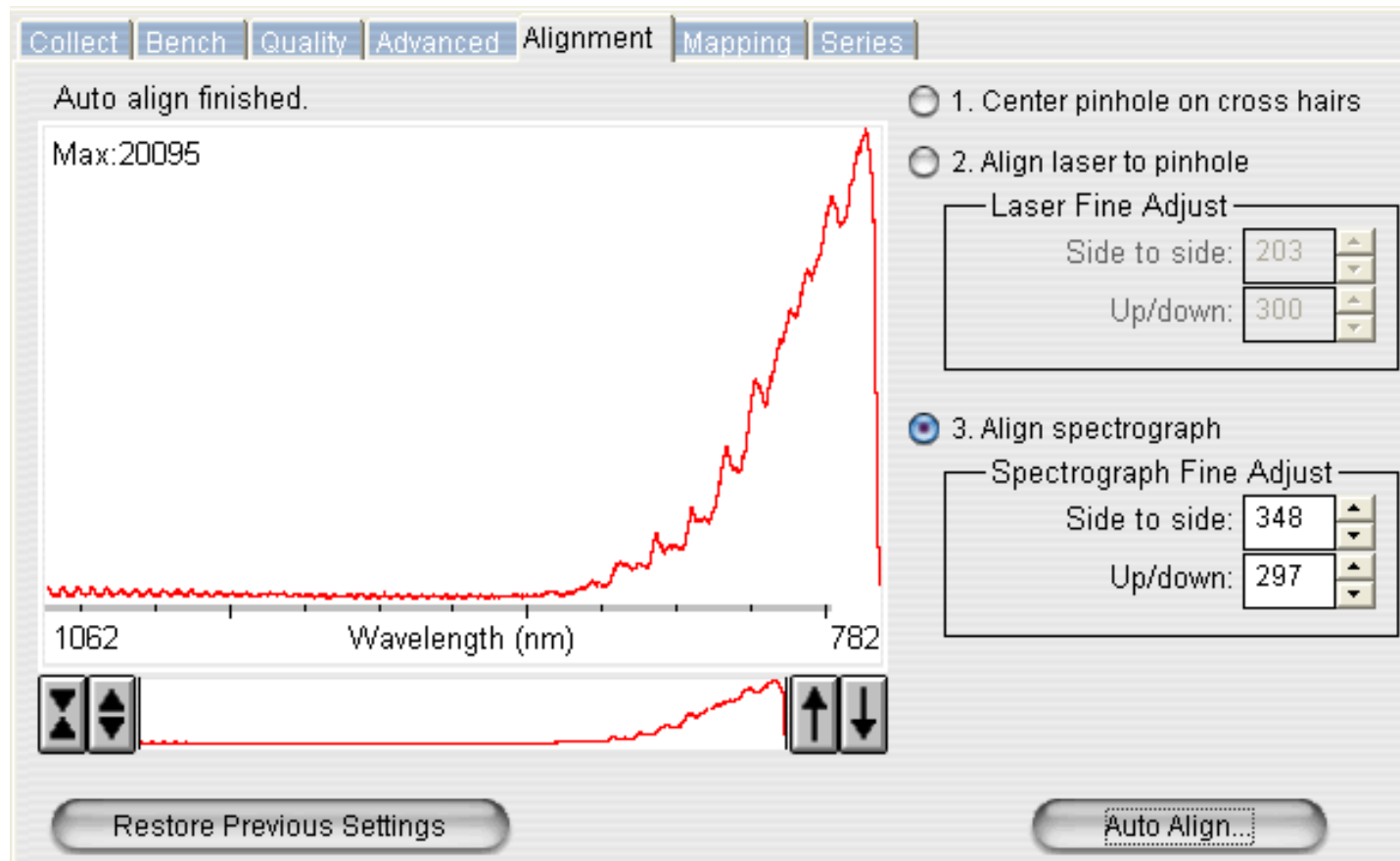
Side to side: 316

Up/down: 290

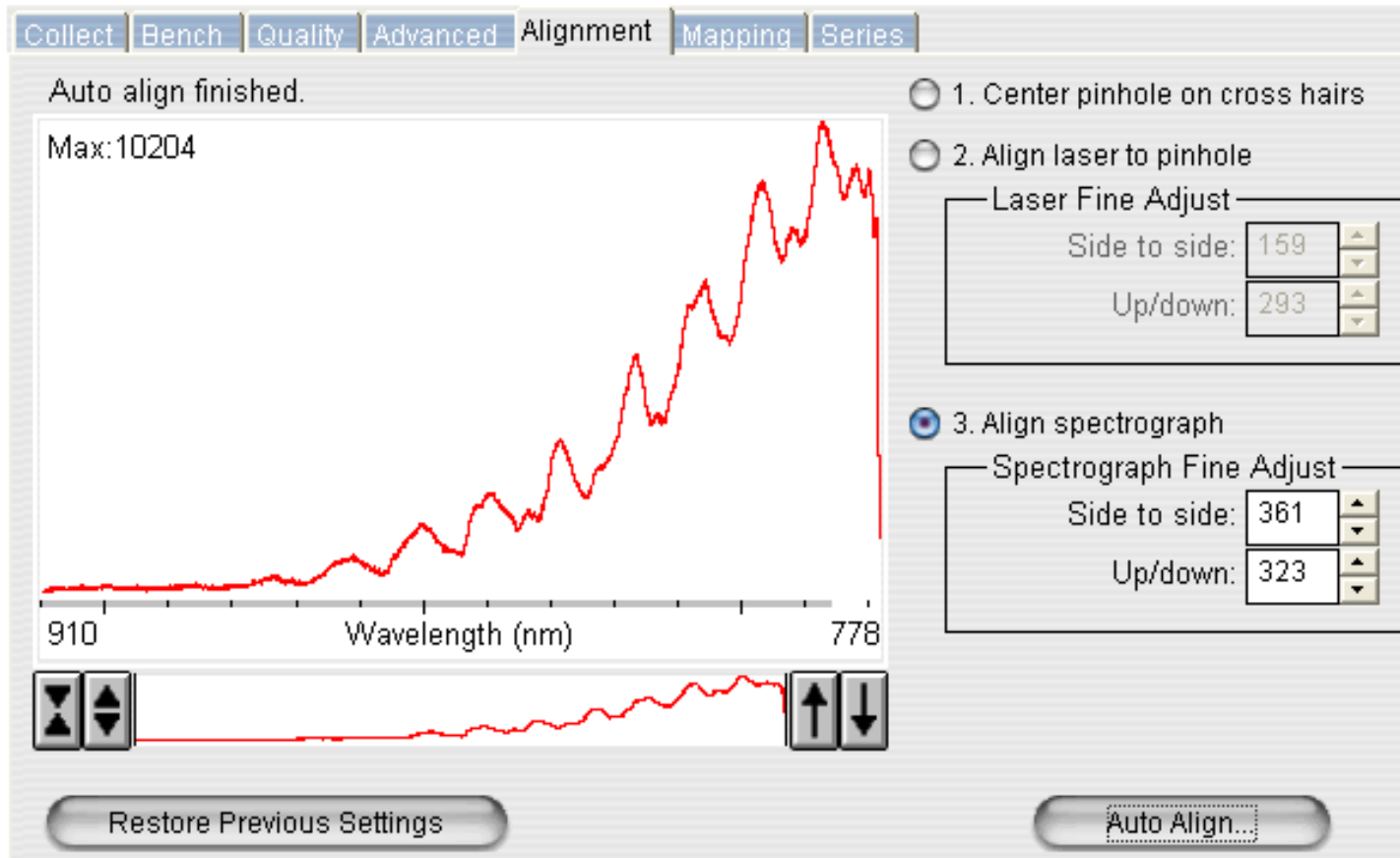
Restore Previous Settings Cancel

# Alignment Tab Microscope (780 full range grating)

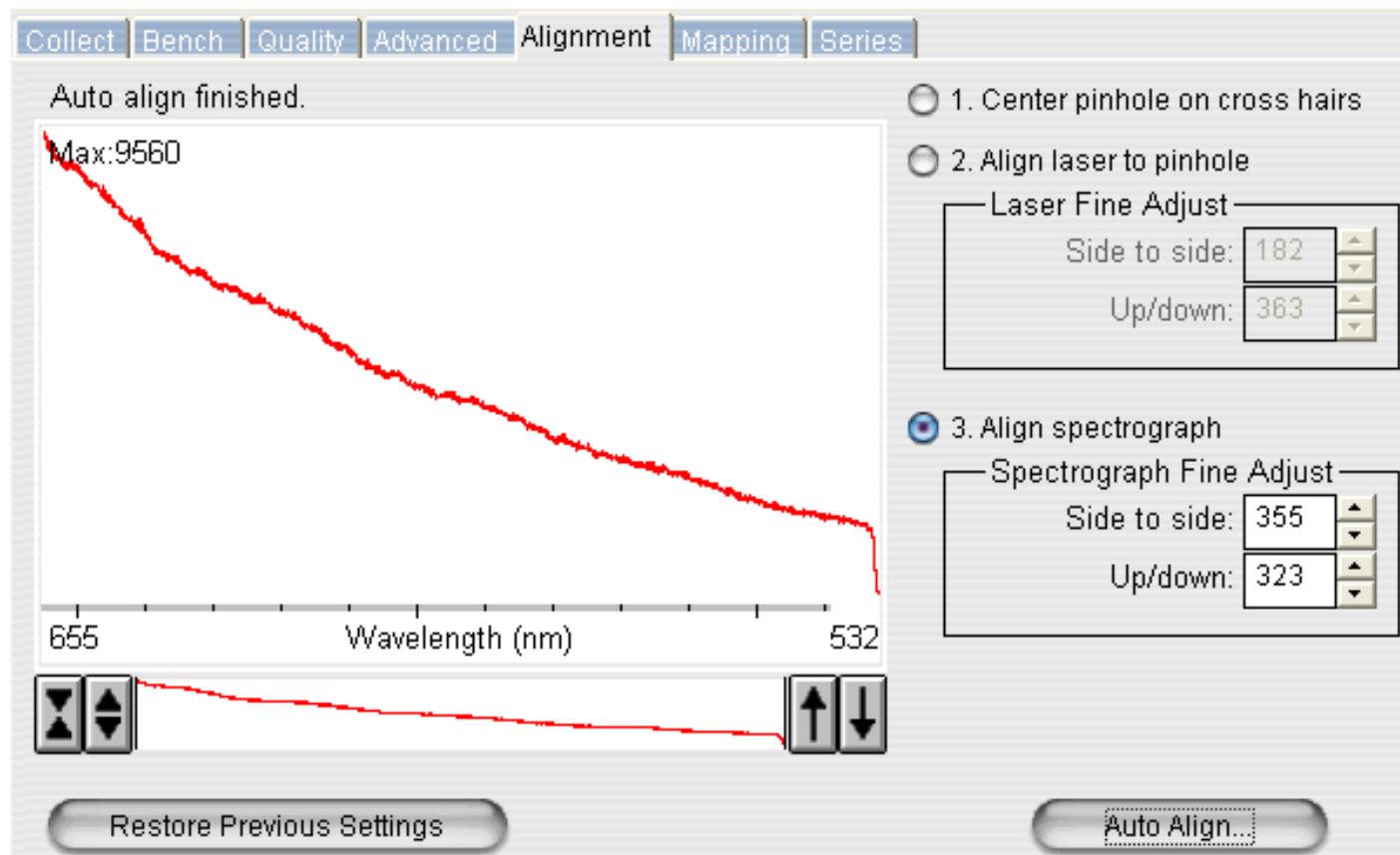
- Step 3, align spectrograph – white light source under pinhole is turned on
- OMNIC moves spectrograph fine adjust lenses to maximize CCD signal
- 780 response is shown below; other lasers and gratings will look different



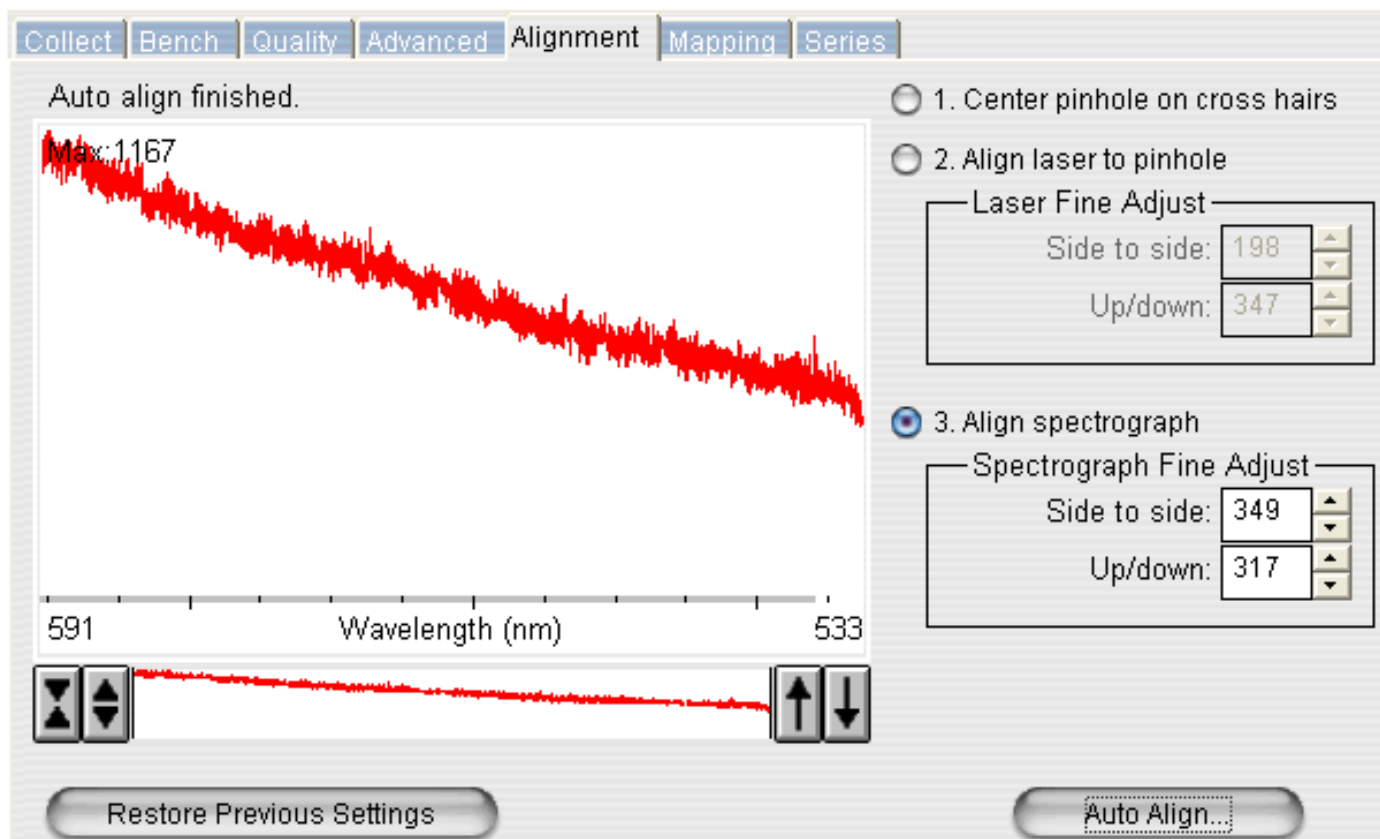
# Alignment Tab Microscope (780 high res grating)



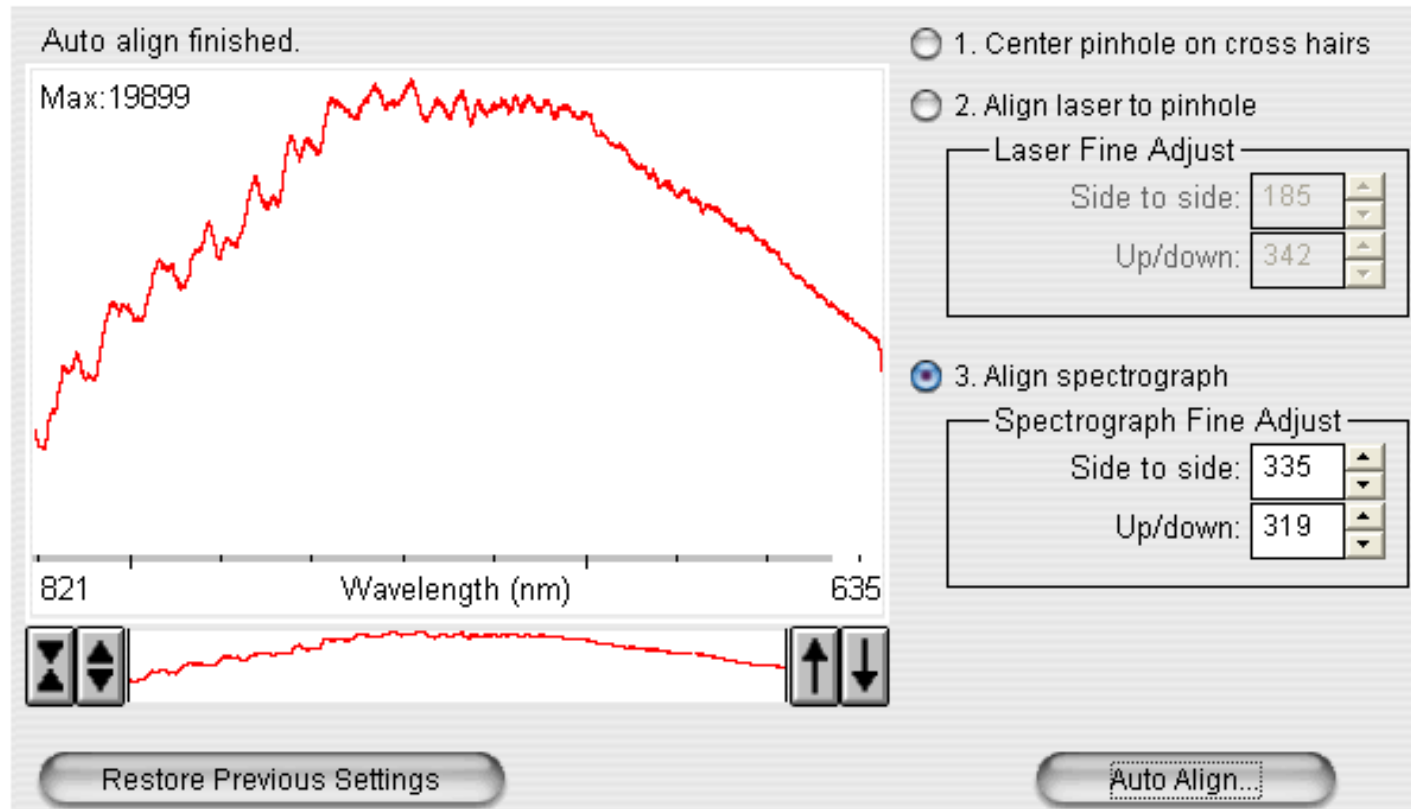
# Alignment Tab Microscope (532 full range grating)



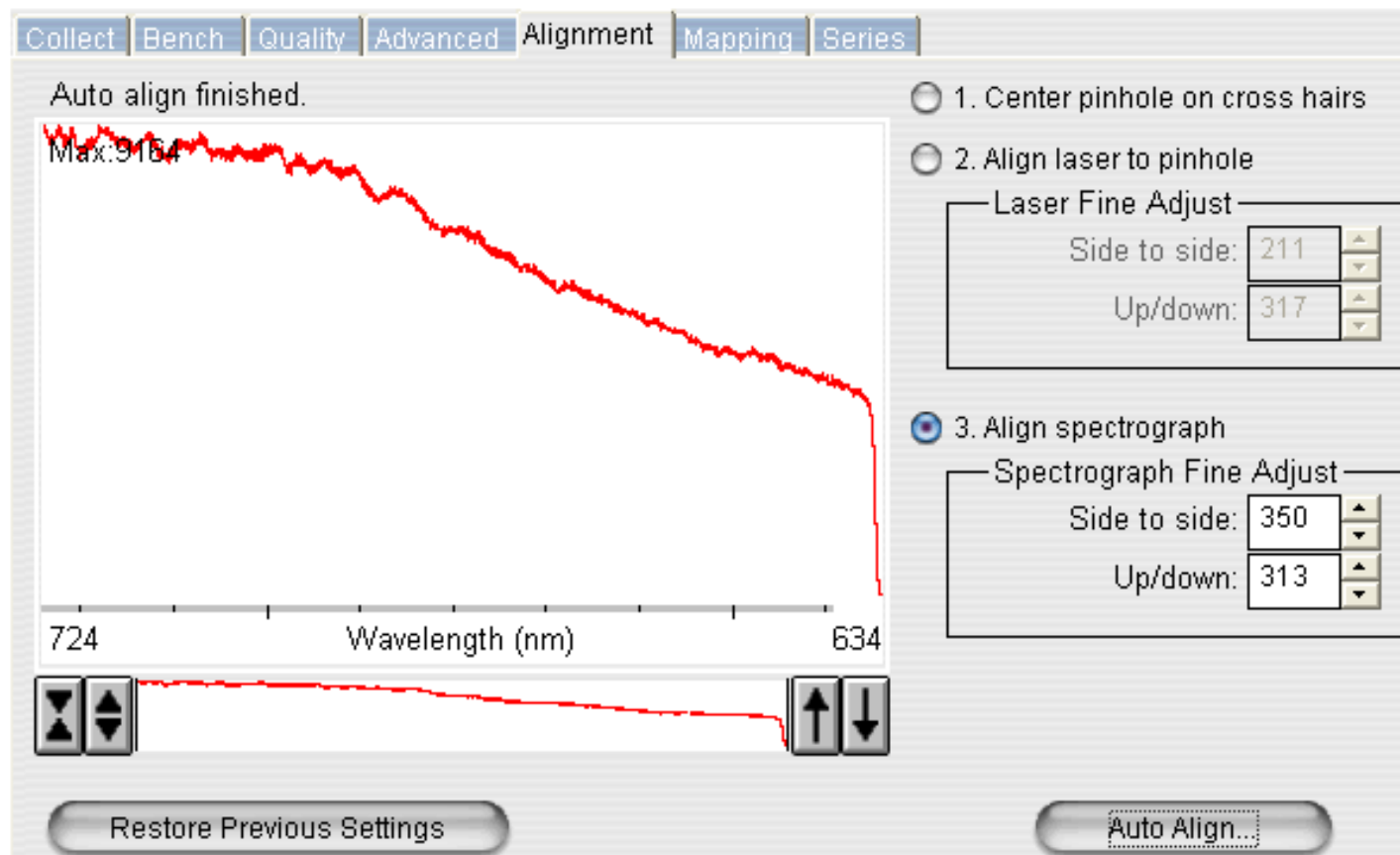
# Alignment Tab Microscope (532 high res grating)



# Alignment Tab Microscope (633 full range grating)



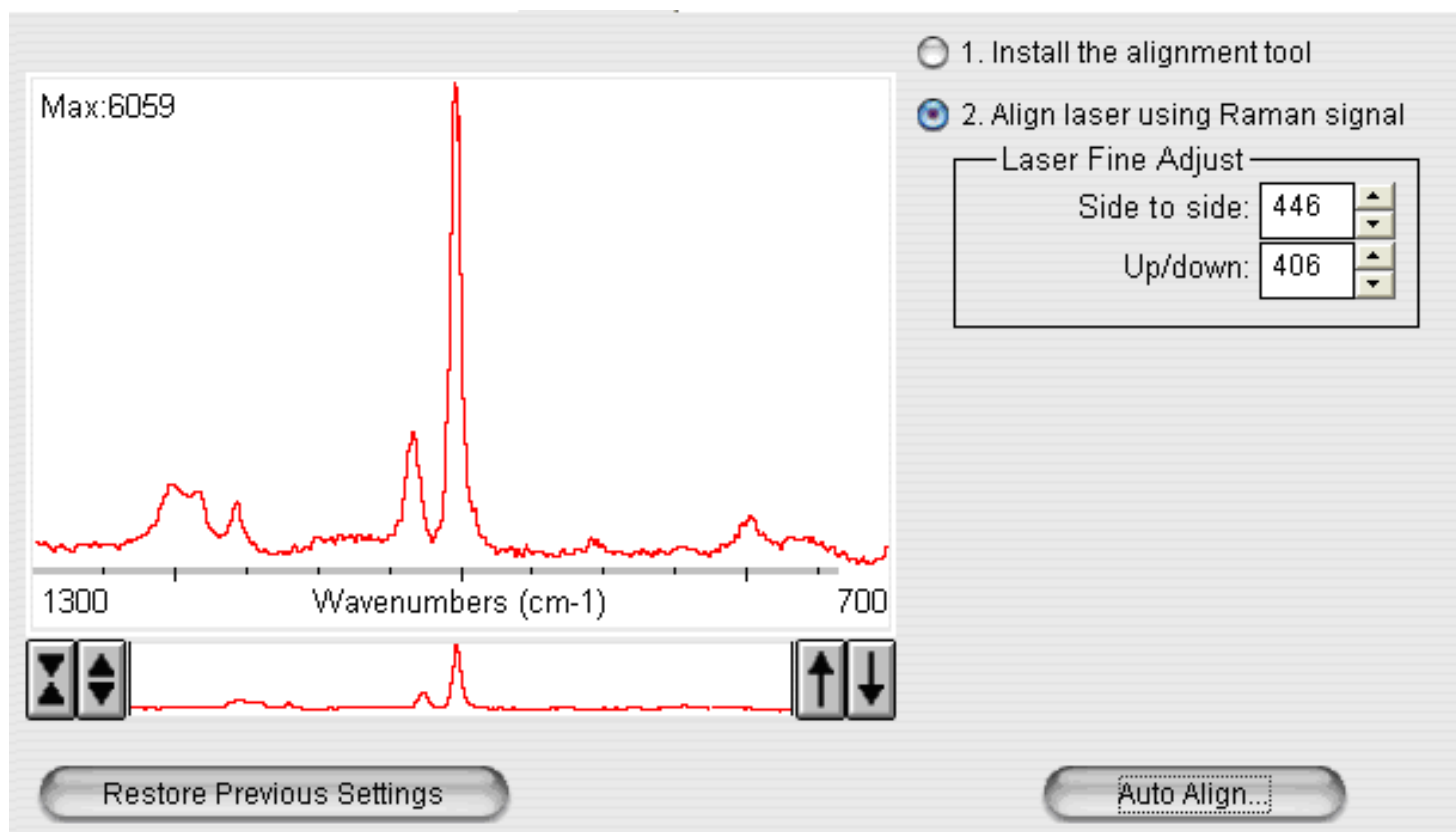
# Alignment Tab Microscope (633 high res grating)





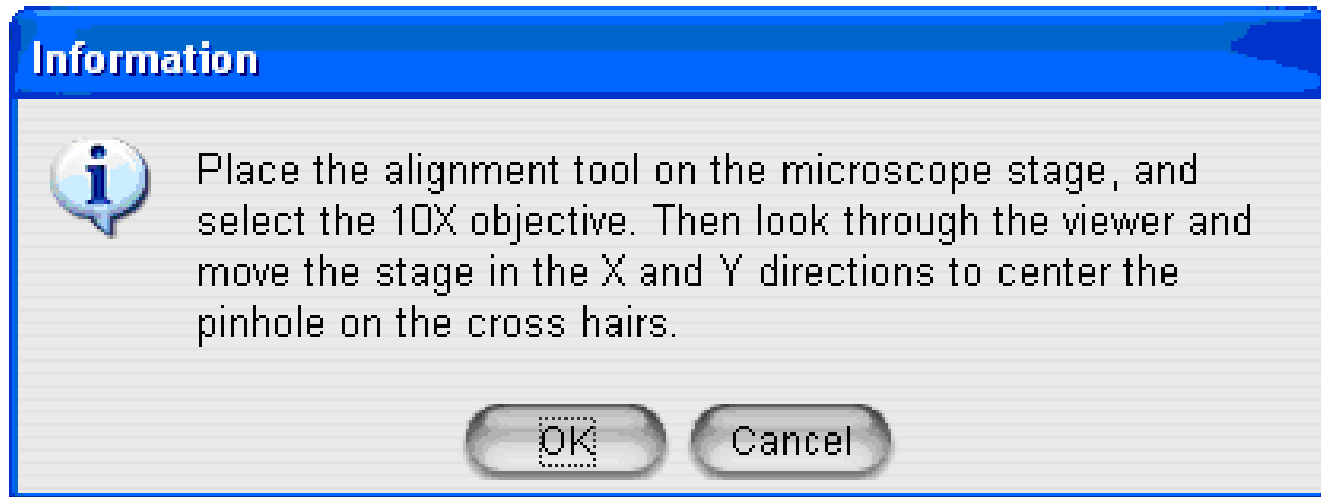
# Alignment Tab DXR SmartRaman

- Align tab will only open if macro align tool is inserted
- Customer clicks button 2
- Polystyrene sample inside align tool moves into beam and the laser fine adjusts maximize the 1001.6  $\text{cm}^{-1}$  peak



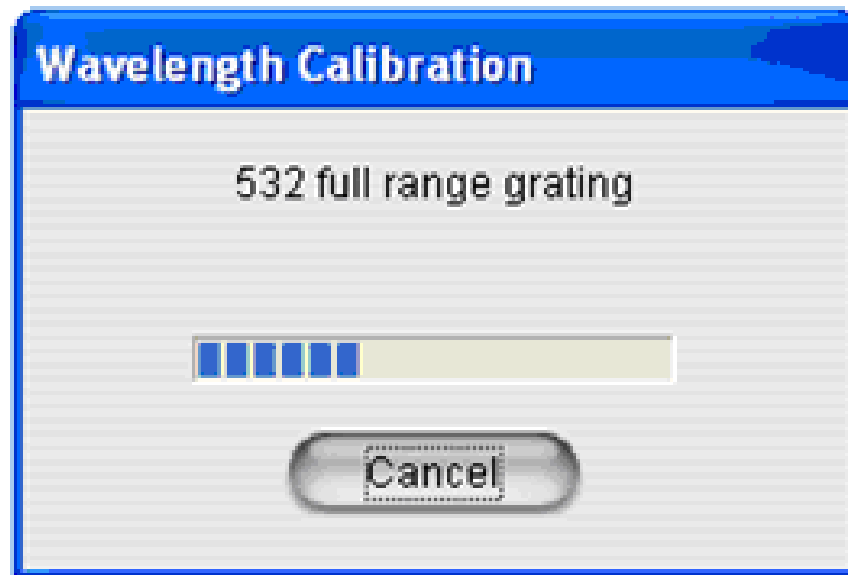
# DXR Microscope Spectrometer Calibration

- Select Calibrate Instrument under Collect menu
- Focus and center the pinhole, then click OK



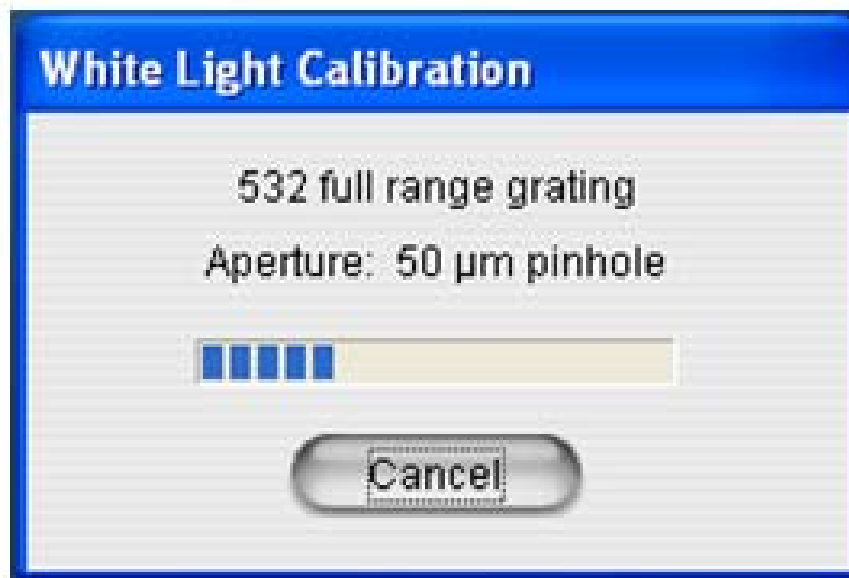
# Spectrometer Calibration - Wavelength

- All three calibrations will be performed automatically with no operator intervention.
- Wavelength calibration moves the neon bulb inside the align tool into the beam and uses the emission lines of the neon bulb to calibrate the x-axis. Multiple neon lines are used so that we can correct for offsets and accordion effects.



# Spectrometer Calibration – White Light

- White light (intensity) calibration is used to remove the instrument function and give more accurate relative band intensities.
- It uses a white light bulb (or red light if the 780 laser is installed).
- The intensity correction is applied automatically to all spectra.



# Spectrometer Calibration – Laser Calibration

- The laser calibration is done by collecting a spectrum from the polystyrene sample inside the alignment tool.
- It compares the measured peak values with the well known polystyrene peak positions and generates a wavenumber correction which is applied to all spectra.



# Spectrometer Calibration – Completion

- All three calibrations must pass.
- If the calibrations fail; reseal the grating and filter, align bench and then rerun calibration.
- Laser calibration in order to be successful, needs to have the laser turned on and needs a successful wavelength calibration.

